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XVII. *Some further Observations on Atmospheric Refraction.**By Stephen Groombridge, Esq. F. R. S.*

Read March 31, 1814.

IN my former paper on atmospherical refraction, communicated to the Royal Society by my late friend, Dr. MASKELYNE, I considered the few observations made below 80° of zenith distance, as not sufficiently to be depended on, for the computation of a general formula of refraction: and I therefore used η Ursæ Majoris ($78^{\circ} 10'$ zen. dis.) as the lowest star for that purpose. Having since applied the computed refraction from the formula thence obtained, to observations of stars below 80° , I have noticed, that such stars so corrected, appeared to be further from the zenith below the Pole, than they ought to have been, from the observations above the Pole: and therefore that the refraction was less at those distances from the zenith, than I had assumed. This has induced me, in the years 1811 and 1812, to make a course of observations of stars below the Pole, above 80° zenith distance; and as near to the horizon, as the trees in Greenwich Park would permit; these being higher than the level of my Observatory. It may also be remarked, that those stars in my former table below 80° , produce the co-latitude in excess; as a confirmation, that the same formula will not apply to those larger arcs, where, from the rapid increase of the tangents, a small error in the assumed quantity becomes more sensible. Although various hypotheses may be formed, from the known density and tem-

perature of the atmosphere; and from these causes may be computed the effect they should have on a ray of light passing through the same: yet we must resort to observation, for the verification of the theory; and reduce the quantity so found, to the most simple and convenient formula. I shall proceed to deduce, from this course of observations, such formulæ as will appear to result, for the computation of the refraction; from the zenith, to the lowest star which I have observed: these may be considered as sufficient for the observation of the sun at the winter solstice, in high latitudes; since those of the moon, from its great parallax, and the planets from their general invisibility, would probably not be attempted. Nevertheless, it is to be wished, as a matter of curiosity, or from which some useful deductions might be made, that in those Observatories, wherein from their elevated situations it might be practicable, the true quantity of refraction should be ascertained to the horizon.

Of all the formulæ for computing the mean refraction, that proposed and used by Dr. BRADLEY, is the most convenient and applicable for the practical astronomer. But as it is now acknowledged, that the numbers he had assumed for the coefficient of r (the refraction;) and of x (the quantity at 45°) were too small: their real values will appear to be the mean of several arcs, and such as I now propose to be adopted. I have found, that the same formula will serve to 87° of zenith distance; possibly this might not happen in low situations, where the height of the vapours would form a greater angle with the horizon: yet in more elevated places, we may reasonably suppose, that a general formula might be carried nearly to the horizon.

In the two annexed tables are given the mean of the observed zenith distances of the several stars; and in the next column the mean refraction, computed from the formula which I proposed in my former paper; viz. $\text{tang. } z - 3.3625 r \times 58''$, 1192: and which has been applied to these observations, corrected for the barometer and thermometer. In the following column, the error or difference is shewn, between the computed and real zenith distances; the assumed mean refraction, corrected by these errors, will give the mean refraction, which should have been applied. From these last quantities are deduced the respective values of (y) the coefficient of r , and of x : and from the mean of the sixteen stars, the resulting numbers will be $y = 3.6342956$, $x = 58.''132967$. Having therefore reduced the whole sixteen by these mean values of y and x ; the error or difference is noted, when compared with the corrected mean refraction which should arise from observation. On a review of these errors, and noticing the mean state of the thermometer for each star, the errors seem to indicate a small correction. I assumed in my former paper, that the refraction as affected by the thermometer, varied in an arithmetical ratio of ,0021 for each degree of FAHRENHEIT'S scale; and the mean state to be at 45° for the thermometer without. Continuing the same mean state, and changing the ratio to ,0020, these errors will be affected $\frac{1}{10000}$ of the refraction, for each degree above and below 45° ; which being applied, will reduce the final error, as shewn in the last column.

In the same manner I proceeded to find the values of y and x , from the six lower stars, contained in the second table; but the respective values of both y and x were variable, and each

in a decreasing ratio. To discover the law of variation for each, would have been complex; therefore retaining the value of x as general, I found y to vary as the minutes of each degree above $87^{\circ} \times .00462$; the co-efficient (y) of r when so reduced is given; the mean refraction resulting, is contained in the following column; and being corrected for the new factors of the thermometer, there remains the final error.

With a view to assist me in ascertaining whether the refractions were affected by local vapours, Dr. BRINKLEY has kindly communicated to me some observations of low stars; which when reduced by the same formula do not materially differ from my own. 12 Can. Ven. at $87^{\circ} 2'$ and α Lyræ at $87^{\circ} 42'$ of which there are the greater number of observations, the former gives the same result within half a second, and the latter $1\frac{1}{2}''$.

Several of the fixed Observatories in Europe being situated in sufficiently high latitudes to obtain the elevation of the pole with much correctness; we are thence enabled, by the circumpolar stars, to find the true quantity of refraction, for all zenith distances: and this having been so ascertained, we may apply the same to the observed zenith distance of the sun, at the winter solstice, as a test of its accuracy. With the smaller quantities of refraction, which were used by Dr. BRADLEY and others, fifty years since, it was not possible, that the latitude deduced from the elevation of the pole, and the mean of the solstices, could agree; the distance of the pole from the equator, so computed, would be less than 90° . Hence also, the two solstices would shew an error of double that difference in the obliquity of the ecliptic, when obtained from the greatest and least zenith distances of the sun. The small number of

observations I have been able to make of the sun, to ascertain the agreement of the two solstices, are subjoined; these appear to confirm the refraction, as deduced from the circum-polar stars.

In the course of my observations I have noticed, that applying the correction for the thermometer without, gives the most accurate result. The difference is very sensible in great zenith distances, from the greater quantity of refraction; and we may reasonably infer, that when the front shutter of the Observatory is opened, the horizontal current of the air is of the same temperature as without; although from the short time of the shutter being opened, during the observation, it is not indicated by the thermometer in the telescope. I have therefore constantly used the thermometer without, for the correction, when I have opened the front shutter; but on all other occasions, I have applied the correction for the thermometer within. The instrument is protected from the horizontal current of air, when the sloping shutters in the roof only are opened; the front shutters being five feet above the graduated circle.

Having formerly proposed certain factors for the thermometrical correction of the refraction, and now finding them relatively the same for the thermometer within and without, changing the ratio for each degree of FAHRENHEIT $\frac{1}{10000}$; I shall adopt, hereafter, the following formula. Putting h° for the degree of the scale; then for the thermometer within, $\overline{49^{\circ} - h^{\circ}} \times ,0023$ when below the mean; $\overline{49^{\circ} - h^{\circ}} \times ,0022$ when above the mean: and for the thermometer without, $\overline{45^{\circ} - h^{\circ}} \times ,0020$; will produce the respective factors. When h° is less than the mean, these will be positive; when greater

than the mean, negative. These factors of different values will be nearly the same for the thermometer within and without, in summer; when the temperatures approximate: but in winter, when the temperatures may differ 6° or 8° , the factors will vary accordingly. It may therefore not be material, when the quantity of refraction is small, not exceeding $1\frac{1}{2}''$, whether the correction is applied for the thermometer within or without; but when the temperature within is supposed to be affected by the horizontal current from without, I should recommend, in all such cases, the use of the latter correction.

To make a direct comparison of these refractions with the French tables, it may be objected, that the latter being computed for the metrical barometer 0,760 m., which is equal to 29,93 inches of our barometer, the mean refraction from the proposed formula should be increased by the factor $+ ,0111$; since we reckon the mean state at 29,60 inches: but, the metrical thermometer at $+ 10$ being equal to ours at 50° ; and the mean state being determined from observation, to be 45° of FAHRENHEIT; the factor for 50° will be $- ,0100$; which will nearly compensate the factor for the barometer. I have therefore compared the mean refractions resulting from my formulæ, with those of the French tables, from 80° to 90° : excepting the two last, the difference is not considerable; and whether these arise from defect in the formulæ, or from local causes, can only be determined by observation.

On comparing the refractions, now proposed, with my former deductions, they will not differ one second at 81° . At 72° the present is $\frac{1}{160}$ of a second less than the former; and at $74^{\circ} 50'$ the zenith distance of the winter solstice, $0,18''$ must be deducted as a correction. It will appear on inspection of

the second table, that β Persei at $88^\circ 1'$ requires but a small equation ($-2''_{.41}$) when computed from my former numbers; and at 88° both these formulæ will coincide. I had therefore, in my former paper, determined too hastily, that the formula proposed would agree with observation so far as 88° ; not having then discovered the discrepancies between 80° and 88° , which are corrected by the present formula.

In order to facilitate the computation of the true refraction, it is required to form a table of mean refractions from certain formulæ. This may appear difficult at the first view; since every refraction must become equal to, $\text{tang. } \overline{z - yr} \times x$; and unless r is assumed very nearly, several operations may be necessary before the differences will vanish. However, proceeding in small arcs of $10'$ to 70° , of $5'$ to 86° , of $4'$ to 88° , of $3'$ to 89° , and of $2'$ thence to the horizon, the second differences of the variation of these arcs may always be taken by inspection: and the resulting refraction will be equal to that assumed, in one operation. The correction for the barometer and thermometer will be the sum of the two factors in the annexed tables, into the mean refraction: and the product added thereto, according to the algebraic sign of their sum, will give the true refraction. This method is more expeditious, than by logarithms; no other tables of reference being required: and the computation will be effected with a small number of figures; which is an object I have constantly had in view.

S. GROOMBRIDGE.

Blackheath, 31 January, 1814.

Reduction of the Solstices.

1810.				Zen. dis. corrected.				1812.				Zen. dis. corrected.				
Dec. 14	-	74. 55. 53.96						June 7	-	28. 0. 25.76						
26	-	52.53						14	-	23.05						
27	-	49.75						15	-	23.74						
30	-	52.66						29	-	23.04						
				<hr/>				July 5				<hr/>				
				74. 55. 52.23								28. 0. 23.80				
Nutation	+ 9.60	} +	1.11					Nutation	- 8.56	} -						
Parallax	- 8.59							Parallax	- 4.04							
Sun's lat.	+ 0.28							Sun's lat.	- 1.01							
Corr. refr.	- 0.18															
				<hr/>								<hr/>				
Latitude				74. 55. 53.34				Latitude				28. 0. 10.19				
				<hr/>								<hr/>				
				51. 28. 2.18								51. 28. 2.18				
				<hr/>								<hr/>				
Mean obliq. eclip.				23. 27. 51.16				Mean obliq. eclip.				23. 27. 51.99				
				<hr/>								<hr/>				
1811.								1812.								
Dec. 9	-	74. 55. 53.64						Dec. 7	-	74. 55. 57.51						
16	-	52.11						8	-	50.51						
22	-	49.79						11	-	55.93						
23	-	55.20						13	-	54.60						
25	-	50.96													<hr/>	
1812 30	-	52.30													74. 55. 54.64	
Jan. 2	-	55.41						Nutation	+ 7.68	} -						
				<hr/>				Parallax	- 8.59							
Nutation	+ 9.16	} +	0.22					Sun's lat.	- 0.52							
Parallax	- 8.59							Corr. refr.	- 0.18							
Sun's lat.	- 0.17											<hr/>				
Corr. refr.	- 0.18											74. 55. 53.03				
				<hr/>								<hr/>				
Latitude				74. 55. 52.99				Latitude				51. 28. 2.18				
				<hr/>								<hr/>				
				51. 28. 2.18								51. 28. 2.18				
				<hr/>								<hr/>				
Mean obliq. eclip.				23. 27. 50.81				Mean obliq. eclip.				23. 27. 50.85				
				<hr/>								<hr/>				

TABLE I.

Star.	No. Obs.	Observed Zenith Distance.	Mean Refraction.	Error from Observ.	Mean Refrac. correct.	Therm.	Refraction at 45°.	Tang. $\frac{z-yr}{x}$.	Error.	Therm. correct.	Final Error.
α Herculis	15	81	39	30.08	6	19.96	6	18.84	6	18.84	0.28
β Capella	30	82	37	49.50	7	6.07	7	4.49	7	4.49	0.26
γ Ursæ maj.	14	82	53	1.76	7	19.80	7	18.11	7	18.11	0.26
β Aurigæ	33	83	29	36.28	7	56.66	7	54.45	7	54.45	0.56
α Cygni	29	83	47	18.26	8	16.55	8	14.09	8	14.09	0.10
α Ursæ maj.	10	84	12	4.14	8	47.17	8	44.21	8	44.21	0.44
α Persei	11	84	15	26.81	8	51.62	8	48.58	8	48.58	0.42
λ Ursæ maj.	10	84	31	48.93	9	14.14	9	10.72	9	10.72	0.38
ϵ Aurigæ	13	84	50	58.25	9	42.81	9	38.85	9	38.85	0.51
ι Lacertæ	15	85	4	10.55	10	4.18	10	0.40	10	0.40	0.50
α Andromedæ	11	85	4	32.79	10	4.79	10	0.40	10	0.40	0.09
ξ Cygni	14	85	10	51.26	10	15.52	10	10.91	10	10.91	0.15
μ Ursæ maj.	12	85	53	46.42	11	38.53	11	31.95	11	31.95	0.18
ι Andromedæ	6	86	6	22.45	12	6.71	12	59.38	12	59.38	0.06
γ Andromedæ	14	86	53	26.64	14	11.52	14	0.29	14	0.29	0.03
θ Andromedæ	6	86	58	29.91	14	27.05	14	15.27	14	15.27	0.59

TABLE II.

β Bootis	10	87	8	19.70	14	58.62	14	47.49	14	47.49	0.19
α Aurigæ	11	87	19	9.83	15	35.64	15	25.91	15	25.91	1.09
γ Aurigæ	15	87	29	7.86	16	11.89	16	3.85	16	3.85	0.70
δ Persei	10	88	1	0.03	18	23.73	18	24.77	18	24.77	0.37
α Cygni	5	88	31	1.12	20	53.78	20	11.37	20	11.37	0.83
γ Persei	5	88	42	35.53	21	59.32	21	26.48	21	26.48	1.08

Mean Refraction.

Zen. dis.	French tables.	S. G.	Value of y .
80	5 19.8	5 19.18	3.63429
81	5 53.5	5 52.83	—
82	6 34.4	6 33.79	—
83	7 24.7	7 24.63	—
84	8 29.9	8 29.13	—
85	9 54.3	9 53.03	—
86	11 48.3	11 45.27	—
87	14 28.1	14 19.80	3.63429
88	18 22.2	18 19.83	3.35709
89	24 21.2	24 32.94	3.07989
90	33 46.3	34 28.13	2.80269

The two formulæ compared.

Zen. dis.	3.36257 × 58".1192	3.634297 × 58".132967
70	2 38.41	2 38.34
71	2 47.31	2 47.23
72	2 57.13	2 57.03
73	3 8.04	3 7.92
74	3 20.22	3 20.07
75	3 33.92	3 33.73
76	3 49.44	3 49.21
77	4 7.19	4 6.89
78	4 27.68	4 27.30
79	4 51.59	4 51.09
80	5 19.85	5 19.18
81	5 53.74	5 52.83
82	6 35.06	6 33.79
83	7 26.46	7 24.63
84	8 31.85	8 29.13
85	9 57.27	9 53.03
86	11 52.21	11 45.27
87	14 31.75	14 19.80
88	18 19.19	18 19.83

Mean Refraction calculated at 58'', 132967 for 45 Degrees of apparent Zenith Distance

Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.
0	0	0.00	9	0	0.20	18	0	0.18.86	27	0	0.29.58	36	0	0.42.17	45	0	0.58.01
10	0	0.17	10	0	0.37	19	0	0.19.05	28	0	0.29.80	37	0	0.42.43	46	0	0.58.35
20	0	0.34	20	0	0.54	20	0	0.19.24	29	0	0.30.00	38	0	0.42.69	47	0	0.58.69
30	0	0.51	30	0	0.72	30	0	0.19.43	30	0	0.30.22	39	0	0.42.95	48	0	0.59.03
40	0	0.68	40	0	0.89	40	0	0.19.61	40	0	0.30.44	40	0	0.43.21	49	0	0.59.38
50	0	0.84	50	0	1.06	50	0	0.19.80	50	0	0.30.65	50	0	0.43.47	50	0	0.59.72
1	0	1.01	10	0	1.24	19	0	0.19.99	28	0	0.30.87	37	0	0.43.74	46	0	1.0.07
10	0	1.18	10	0	1.04	10	0	0.20.18	10	0	0.31.09	10	0	0.44.00	10	0	1.0.42
20	0	1.35	20	0	1.09	20	0	0.20.37	20	0	0.31.30	20	0	0.44.27	20	0	1.0.77
30	0	1.52	30	0	1.07	30	0	0.20.56	30	0	0.31.52	30	0	0.44.53	30	0	1.1.13
40	0	1.69	40	0	1.04	40	0	0.20.75	40	0	0.31.74	40	0	0.44.80	40	0	1.1.48
50	0	1.86	50	0	1.11	50	0	0.20.94	50	0	0.31.96	50	0	0.45.07	50	0	1.1.84
2	0	2.03	11	0	1.12	20	0	0.21.13	29	0	0.32.18	38	0	0.45.34	47	0	1.2.20
10	0	2.20	10	0	1.16	10	0	0.21.32	10	0	0.32.40	10	0	0.45.62	10	0	1.2.57
20	0	2.37	20	0	1.14	20	0	0.21.51	20	0	0.32.62	20	0	0.45.89	20	0	1.2.93
30	0	2.53	30	0	1.18	30	0	0.21.71	30	0	0.32.85	30	0	0.46.16	30	0	1.3.30
40	0	2.70	40	0	1.19	40	0	0.21.90	40	0	0.33.07	40	0	0.46.44	40	0	1.3.67
50	0	2.87	50	0	1.27	50	0	0.22.09	50	0	0.33.29	50	0	0.46.72	50	0	1.4.04
3	0	3.04	12	0	1.23	21	0	0.22.29	30	0	0.33.52	39	0	0.47.00	48	0	1.4.41
10	0	3.21	10	0	1.25	10	0	0.22.48	10	0	0.33.74	10	0	0.47.28	10	0	1.4.79
20	0	3.38	20	0	1.27	20	0	0.22.67	20	0	0.33.97	20	0	0.47.56	20	0	1.5.17
30	0	3.55	30	0	1.28	30	0	0.22.87	30	0	0.34.20	30	0	0.47.84	30	0	1.5.55
40	0	3.72	40	0	1.30	40	0	0.23.06	40	0	0.34.42	40	0	0.48.12	40	0	1.5.94
50	0	3.89	50	0	1.32	50	0	0.23.26	50	0	0.34.65	50	0	0.48.41	50	0	1.6.33
4	0	4.06	13	0	1.34	22	0	0.23.46	31	0	0.34.88	40	0	0.48.69	49	0	1.6.72
10	0	4.23	10	0	1.35	10	0	0.23.63	10	0	0.35.11	10	0	0.48.98	10	0	1.7.11
20	0	4.40	20	0	1.37	20	0	0.23.85	20	0	0.35.34	20	0	0.49.27	20	0	1.7.50
30	0	4.57	30	0	1.39	30	0	0.24.05	30	0	0.35.57	30	0	0.49.56	30	0	1.7.90
40	0	4.74	40	0	1.41	40	0	0.24.25	40	0	0.35.81	40	0	0.49.85	40	0	1.8.30
50	0	4.91	50	0	1.43	50	0	0.24.45	50	0	0.36.04	50	0	0.50.15	50	0	1.8.70
5	0	5.08	14	0	1.44	23	0	0.24.65	32	0	0.36.27	41	0	0.50.44	50	0	1.9.11
10	0	5.25	10	0	1.46	10	0	0.24.85	10	0	0.36.51	10	0	0.50.74	10	0	1.9.52
20	0	5.42	20	0	1.48	20	0	0.25.05	20	0	0.36.74	20	0	0.51.04	20	0	1.9.93
30	0	5.59	30	0	1.50	30	0	0.25.25	30	0	0.36.98	30	0	0.51.34	30	0	1.10.34
40	0	5.76	40	0	1.52	40	0	0.25.45	40	0	0.37.22	40	0	0.51.64	40	0	1.10.76
50	0	5.93	50	0	1.53	50	0	0.25.65	50	0	0.37.46	50	0	0.51.94	50	0	1.11.18
6	0	6.10	15	0	1.55	24	0	0.25.85	33	0	0.37.70	42	0	0.52.25	51	0	1.11.60
10	0	6.27	10	0	1.57	10	0	0.26.05	10	0	0.37.94	10	0	0.52.55	10	0	1.12.03
20	0	6.44	20	0	1.59	20	0	0.26.26	20	0	0.38.18	20	0	0.52.86	20	0	1.12.46
30	0	6.62	30	0	1.61	30	0	0.26.46	30	0	0.38.42	30	0	0.53.17	30	0	1.12.89
40	0	6.79	40	0	1.62	40	0	0.26.66	40	0	0.38.66	40	0	0.53.48	40	0	1.13.33
50	0	6.96	50	0	1.64	50	0	0.26.87	50	0	0.38.91	50	0	0.53.79	50	0	1.13.77
7	0	7.13	16	0	1.66	25	0	0.27.07	34	0	0.39.15	43	0	0.54.11	52	0	1.14.21
10	0	7.30	10	0	1.68	10	0	0.27.28	10	0	0.39.40	10	0	0.54.42	10	0	1.14.65
20	0	7.47	20	0	1.70	20	0	0.27.49	20	0	0.39.64	20	0	0.54.74	20	0	1.15.10
30	0	7.64	30	0	1.72	30	0	0.27.69	30	0	0.39.89	30	0	0.55.06	30	0	1.15.55
40	0	7.82	40	0	1.73	40	0	0.27.90	40	0	0.40.14	40	0	0.55.38	40	0	1.16.01
50	0	7.99	50	0	1.75	50	0	0.28.11	50	0	0.40.39	50	0	0.55.70	50	0	1.16.47
8	0	8.16	17	0	1.77	26	0	0.28.32	35	0	0.40.64	44	0	0.56.03	53	0	1.16.93
10	0	8.33	10	0	1.79	10	0	0.28.53	10	0	0.40.89	10	0	0.56.35	10	0	1.17.39
20	0	8.50	20	0	1.81	20	0	0.28.74	20	0	0.41.15	20	0	0.56.68	20	0	1.17.86
30	0	8.68	30	0	1.83	30	0	0.28.95	30	0	0.41.42	30	0	0.57.01	30	0	1.18.33
40	0	8.85	40	0	1.84	40	0	0.29.16	40	0	0.41.66	40	0	0.57.34	40	0	1.18.81
50	0	9.02	50	0	1.86	50	0	0.29.37	50	0	0.41.91	50	0	0.57.68	50	0	1.19.29

7 for 45 Degrees of apparent Zenith Distance \times Tang. $Z = 3.6342956r$ so far as 87° Zen. Dis. below wh

[illegible]

3.6342956r so far as 87° Zen. Dis. below which *r* is reduced, 00462 for each minute.

[To face page 347.]

Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.	Zen. Dis.	Refrac.	Diff.
3 0	1 53.53	0.82	71 0	2 47.23	0.78	75 0	3 41.22	1.30	80 0	5 19.18	2.57	84 0	9 8.18	7.05	88 0	15 54.47	17.05
10	1 54.35	0.82	5	2 48.01	0.79	35	3 42.52	1.31	5	5 21.75	2.61	35	9 15.23	7.20	21	20 11.53	17.45
20	1 55.17	0.83	10	2 48.80	0.79	40	3 43.83	1.32	10	5 24.36	2.65	40	9 22.43	7.38	24	20 28.98	17.84
30	1 56.00	0.85	15	2 49.59	0.80	45	3 45.15	1.34	15	5 27.01	2.70	45	9 29.81	7.56	27	20 46.82	18.26
40	1 56.85	0.85	20	2 50.39	0.81	50	3 46.49	1.35	20	5 29.71	2.73	50	9 37.37	7.73	30	21 5.08	18.68
50	1 57.70	0.86	25	2 51.20	0.81	55	3 47.84	1.37	25	5 32.44	2.77	55	9 45.10	7.93	33	21 23.76	19.12
4 0	1 58.56	0.87	30	2 52.01	0.82	76 0	3 49.21	1.38	30	5 35.21	2.82	85 0	9 53.03	8.12	36	21 42.88	19.57
10	1 59.43	0.89	35	2 52.83	0.82	5	3 50.59	1.40	35	5 38.03	2.87	5	10 1.15	8.31	39	22 2.45	20.02
20	2 0.32	0.89	40	2 53.65	0.84	10	3 51.99	1.42	40	5 40.90	2.91	10	10 9.46	8.53	42	22 22.47	20.49
30	2 1.21	0.91	45	2 54.49	0.84	15	3 53.41	1.43	45	5 43.81	2.96	15	10 17.99	8.74	45	22 42.96	20.99
40	2 2.12	0.91	50	2 55.33	0.85	20	3 54.84	1.45	50	5 46.77	3.01	20	10 26.73	8.96	48	23 3.95	21.47
50	2 3.03	0.93	55	2 56.18	0.85	25	3 56.29	1.46	55	5 49.78	3.05	25	10 35.69	9.19	51	23 25.42	21.98
5 0	2 3.96	0.94	72 0	2 57.03	0.86	30	3 57.75	1.48	81 0	5 52.83	3.10	30	10 44.88	9.43	54	23 47.40	22.50
10	2 4.90	0.95	5	2 57.89	0.87	35	3 59.23	1.50	5	5 55.93	3.16	35	10 54.31	9.67	57	24 9.90	23.04
20	2 5.85	0.96	10	2 58.76	0.88	40	4 0.73	1.52	10	5 59.09	3.21	40	11 3.98	9.92	89 0	24 32.94	23.44
30	2 6.81	0.97	15	2 59.64	0.89	45	4 2.25	1.53	15	6 2.30	3.26	45	11 13.90	10.19	2	24 48.60	23.84
40	2 7.78	0.99	20	3 0.53	0.90	50	4 3.78	1.55	20	6 5.56	3.32	50	11 24.09	10.45	4	25 4.51	24.24
50	2 8.77	1.00	25	3 1.43	0.90	55	4 5.33	1.56	25	6 8.88	3.38	55	11 34.54	10.73	6	25 20.67	24.64
6 0	2 9.77	1.01	30	3 2.33	0.91	77 0	4 6.89	1.59	30	6 12.26	3.43	86 0	11 45.27	8.80	8	25 37.09	25.04
10	2 10.78	1.03	35	3 3.24	0.92	5	4 8.48	1.61	35	6 15.69	3.50	4	11 54.07	8.99	10	25 53.77	25.44
20	2 11.81	1.04	40	3 4.16	0.93	10	4 10.09	1.63	40	6 19.19	3.55	8	12 2.06	9.18	12	26 10.71	25.84
30	2 12.85	1.05	45	3 5.09	0.93	15	4 11.72	1.65	45	6 22.74	3.62	12	12 12.24	9.39	14	26 27.91	26.24
40	2 13.90	1.07	50	3 6.02	0.95	20	4 13.37	1.66	50	6 26.36	3.68	16	12 21.63	9.59	16	26 45.40	26.64
50	2 14.97	1.08	55	3 6.97	0.95	25	4 15.03	1.69	55	6 30.04	3.75	20	12 31.22	9.80	18	27 3.16	27.04
7 0	2 16.05	1.09	73 0	3 7.92	0.96	30	4 16.72	1.71	82 0	6 33.79	3.82	24	12 41.02	10.02	20	27 21.20	27.44
10	2 17.14	1.11	5	3 8.88	0.97	35	4 18.43	1.73	5	6 37.61	3.89	28	12 51.04	10.24	22	27 39.52	27.84
20	2 18.25	1.13	10	3 9.85	0.98	40	4 20.16	1.75	10	6 41.50	3.96	32	13 1.28	10.47	24	27 58.14	28.24
30	2 19.38	1.14	15	3 10.83	0.99	45	4 21.91	1.77	15	6 45.46	4.03	36	13 11.75	10.71	26	28 17.05	28.64
40	2 20.52	1.16	20	3 11.82	1.00	50	4 23.68	1.80	20	6 49.49	4.10	40	13 22.46	10.96	28	28 36.26	29.04
50	2 21.68	1.17	25	3 12.82	1.00	55	4 25.48	1.82	25	6 53.59	4.19	44	13 33.42	11.20	30	28 55.78	29.44
8 0	2 22.85	1.19	30	3 13.82	1.02	78 0	4 27.30	1.84	30	6 57.78	4.26	48	13 44.62	11.46	32	29 15.60	29.84
10	2 24.04	1.21	35	3 14.84	1.02	5	4 29.14	1.87	35	7 2.04	4.35	52	13 56.08	11.73	34	29 35.74	30.24
20	2 25.25	1.22	40	3 15.86	1.04	10	4 31.01	1.89	40	7 6.39	4.43	56	14 7.81	11.99	36	29 56.19	30.64
30	2 26.47	1.24	45	3 16.90	1.05	15	4 32.90	1.91	45	7 10.82	4.51	87 0	14 19.80	13.11	38	30 16.97	31.04
40	2 27.71	1.26	50	3 17.95	1.06	20	4 34.81	1.94	50	7 15.33	4.60	4	14 32.91	13.46	40	30 38.07	31.44
50	2 28.97	1.28	55	3 19.01	1.06	25	4 36.75	1.97	55	7 19.93	4.70	8	14 46.37	13.83	42	30 59.51	31.84
9 0	2 30.25	1.30	74 0	3 20.07	1.08	30	4 38.72	1.99	83 0	7 24.63	4.79	12	15 0.20	14.20	44	31 21.28	32.24
10	2 31.55	1.32	5	3 21.15	1.08	35	4 40.71	2.02	5	7 29.42	4.88	16	15 14.40	14.60	46	31 43.39	32.64
20	2 32.87	1.34	10	3 22.23	1.10	40	4 42.73	2.05	10	7 34.30	4.99	20	15 29.00	15.00	48	32 5.85	33.04
30	2 34.21	1.35	15	3 23.33	1.11	45	4 44.78	2.07	15	7 39.29	5.09	24	15 44.00	15.42	50	32 28.66	33.44
40	2 35.56	1.38	20	3 24.44	1.12	50	4 46.85	2.11	20	7 44.38	5.19	28	15 59.42	15.85	52	32 51.82	33.84
50	2 36.94	1.40	25	3 25.56	1.13	55	4 48.96	2.13	25	7 49.57	5.30	32	16 15.27	16.31	54	33 15.35	34.24
10 0	2 38.34	0.71	30	3 26.69	1.15	79 0	4 51.09	2.16	30	7 54.87	5.41	36	16 31.58	16.77	56	33 39.24	34.64
5	2 39.05	0.71	35	3 27.84	1.15	5	4 53.25	2.19	35	8 0.28	5.52	40	16 48.35	17.26	58	34 3.50	35.04
10	2 39.76	0.72	40	3 28.99	1.17	10	4 55.44	2.23	40	8 5.80	5.65	44	17 5.61	17.76	90 0	34 28.13	35.44
15	2 40.48	0.73	45	3 30.16	1.18	15	4 57.67	2.25	45	8 11.45	5.76	48	17 23.37	18.27	2	34 53.15	35.84
20	2 41.21	0.73	50	3 31.34	1.19	20	4 59.92	2.29	50	8 17.21	5.89	52	17 41.64	18.82	4	35 18.55	36.24
25	2 41.94	0.74	55	3 32.53	1.20	25	5 2.21	2.32	55	8 23.10	6.03	56	18 0.46	19.37	6	35 44.34	36.64
30	2 42.68	0.74	75 0	3 33.73	1.22	30	5 4.53	2.35	84 0	8 29.13	6.15	88 0	18 19.83	14.91	8	36 10.52	37.04
35	2 43.42	0.75	5	3 34.95	1.23	35	5 6.88	2.39	5	8 35.28	6.29	3	18 34.74	15.24	10	36 37.10	37.44
40	2 44.17	0.76	10	3 36.18	1.24	40	5 9.27	2.42	10	8 41.57	6.43	6	18 49.98	15.59	12	37 4.08	37.84
45	2 44.93	0.76	15	3 37.42	1.25	45	5 11.69	2.46	15	8 48.00	6.57	9	19 5.57	15.94	14	37 31.47	38.24
50	2 45.69	0.77	20	3 38.67	1.27	50	5 14.15	2.50	20	8 54.57	6.73	12	19 21.51	16.29	16	37 59.28	38.64
55	2 46.46	0.77	25	3 39.94	1.28	55	5 16.65	2.53	25	9 1.30	6.88	15	19 37.80	16.67	18	38 27.50	39.04

Barometer
in inches.

FAHRENHEIT's Thermometer.

Factors for the correction of the mean Refraction.

	-		+
28.60	,0350	29.60	,0000
62	,0342	62	,0007
64	,0335	64	,0014
66	,0328	66	,0020
68	,0321	68	,0027
28.70	,0314	29.70	,0034
72	,0306	72	,0041
74	,0299	74	,0047
76	,0292	76	,0054
78	,0285	78	,0061
28.80	,0278	29.80	,0068
82	,0271	82	,0074
84	,0264	84	,0081
86	,0256	86	,0088
88	,0249	88	,0095
28.90	,0242	29.90	,0101
92	,0235	92	,0108
94	,0228	94	,0115
96	,0221	96	,0122
98	,0214	98	,0128
29.00	,0207	30.00	,0135
02	,0200	02	,0142
04	,0193	04	,0149
06	,0186	06	,0155
08	,0179	08	,0162
29.10	,0172	30.10	,0169
12	,0165	12	,0176
14	,0158	14	,0182
16	,0151	16	,0189
18	,0144	18	,0196
29.20	,0137	30.20	,0203
22	,0130	22	,0210
24	,0123	24	,0216
26	,0116	26	,0223
28	,0109	28	,0230
29.30	,0102	30.30	,0237
32	,0096	32	,0243
34	,0089	34	,0250
36	,0082	36	,0257
38	,0075	38	,0264
29.40	,0068	30.40	,0270
42	,0061	42	,0277
44	,0054	44	,0284
46	,0048	46	,0291
48	,0041	48	,0297
29.50	,0034	30.50	,0304
52	,0027	52	,0311
54	,0020	54	,0318
56	,0014	56	,0324
58	,0007	58	,0331

°	Within. +	Without. +	°	Within. -	Without. -
21.5	,0632	,0470	49.0	,0000	,0080
22.0	,0621	,0460	49.5	,0011	,0090
22.5	,0609	,0450	50.0	,0022	,0100
23.0	,0598	,0440	50.5	,0033	,0110
23.5	,0586	,0430	51.0	,0044	,0120
24.0	,0575	,0420	51.5	,0055	,0130
24.5	,0563	,0410	52.0	,0066	,0140
25.0	,0552	,0400	52.5	,0077	,0150
25.5	,0540	,0390	53.0	,0088	,0160
26.0	,0529	,0380	53.5	,0099	,0170
26.5	,0517	,0370	54.0	,0110	,0180
27.0	,0506	,0360	54.5	,0121	,0190
27.5	,0494	,0350	55.0	,0132	,0200
28.0	,0483	,0340	55.5	,0143	,0210
28.5	,0471	,0330	56.0	,0154	,0220
29.0	,0460	,0320	56.5	,0165	,0230
29.5	,0448	,0310	57.0	,0176	,0240
30.0	,0437	,0300	57.5	,0187	,0250
30.5	,0425	,0290	58.0	,0198	,0260
31.0	,0414	,0280	58.5	,0209	,0270
31.5	,0402	,0270	59.0	,0220	,0280
32.0	,0391	,0260	59.5	,0231	,0290
32.5	,0379	,0250	60.0	,0242	,0300
33.0	,0368	,0240	60.5	,0253	,0310
33.5	,0356	,0230	61.0	,0264	,0320
34.0	,0345	,0220	61.5	,0275	,0330
34.5	,0333	,0210	62.0	,0286	,0340
35.0	,0322	,0200	62.5	,0297	,0350
35.5	,0310	,0190	63.0	,0308	,0360
36.0	,0299	,0180	63.5	,0319	,0370
36.5	,0287	,0170	64.0	,0330	,0380
37.0	,0276	,0160	64.5	,0341	,0390
37.5	,0264	,0150	65.0	,0352	,0400
38.0	,0253	,0140	65.5	,0363	,0410
38.5	,0241	,0130	66.0	,0374	,0420
39.0	,0230	,0120	66.5	,0385	,0430
39.5	,0218	,0110	67.0	,0396	,0440
40.0	,0207	,0100	67.5	,0407	,0450
40.5	,0195	,0090	68.0	,0418	,0460
41.0	,0184	,0080	68.5	,0429	,0470
41.5	,0172	,0070	69.0	,0440	,0480
42.0	,0161	,0060	69.5	,0451	,0490
42.5	,0149	,0050	70.0	,0462	,0500
43.0	,0138	,0040	70.5	,0473	,0510
43.5	,0126	,0030	71.0	,0484	,0520
44.0	,0115	,0020	71.5	,0495	,0530
44.5	,0103	,0010	72.0	,0506	,0540
45.0	,0092	—	72.5	,0517	,0550
45.5	,0080	,0010	73.0	,0528	,0560
46.0	,0069	,0020	73.5	,0539	,0570
46.5	,0057	,0030	74.0	,0550	,0580
47.0	,0046	,0040	74.5	,0561	,0590
47.5	,0034	,0050	75.0	,0572	,0600
48.0	,0023	,0060	75.5	,0583	,0610
48.5	,0011	,0070	76.0	,0594	,0620